

## Research Article

# Unitarity of Singh-Hagen Model in $D$ Dimensions

Elias L. Mendonça  and R. Schmidt Bittencourt

UNESP-Campus de Guaratinguetá-DFQ, Av. Dr. Ariberto Pereira da Cunha, 333, CEP 12516-410 Guaratinguetá, SP, Brazil

Correspondence should be addressed to Elias L. Mendonça; [elias.leite@unesp.br](mailto:elias.leite@unesp.br)

Received 19 November 2019; Revised 27 January 2020; Accepted 8 February 2020; Published 6 March 2020

Academic Editor: Shi-Hai Dong

Copyright © 2020 Elias L. Mendonça and R. Schmidt Bittencourt. This is an open access article distributed under the Creative Commons Attribution License, which permits unrestricted use, distribution, and reproduction in any medium, provided the original work is properly cited. The publication of this article was funded by SCOAP<sup>3</sup>.

The particle content of the Singh-Hagen model (SH) in  $D$  dimensions is revisited. We suggest a complete set of spin-projection operators acting on totally symmetric rank-3 fields. We give a general expression for the propagator and determine the coefficients of the SH model confirming previous results of the literature. Adding source terms, we provide a unitarity analysis in  $D$  dimensions. In addition, we have also analyzed the positivity of the massless Hamiltonian.

## 1. Introduction

The suggestion of a free theory describing higher spin particles dates back to 1936 by Dirac [1] and 1939 by Fierz and Pauli (FP) [2]. As fundamental assumptions, such theories should be invariant under Poincaré transformations and at the same time guarantee the energy positivity. Particularly in the case of higher spin theories, positivity deserves special attention, since attempting to describe such particles, we are faced with extra propagation modes with spins lower than those we would like to describe; such modes may be ghosts in some cases. In order to remove the spurious degrees of freedom, one needs the addition of auxiliary fields, which sometimes makes the analysis of the equations of motion truly complicated.

From the experience with lower spins, one knows that the particle content of some theory may be directly obtained by calculating the propagator, but in order to obtain it, one needs to construct a complete basis of spin-projection operators. Strictly speaking about bosonic examples, it is quite simple to obtain the propagator of a rank-one field theory with the help of the transverse  $\theta_{\mu\nu} = \eta_{\mu\nu} - \omega_{\mu\nu}$  and longitudinal  $\omega_{\mu\nu} = \partial_\mu \partial_\nu / \square$  operators. By mean of these projectors, Barnes and Rivers [3] have introduced a complete set of spin-projection operators which allows us to determine the particle content of a given rank-two field theory (a slightly different basis is also used by [4]). Some extensions of this set of

projectors are given at [5, 6] where a new class of projection operators for three-dimensional models is constructed.

The spin-3 case is the simplest bosonic example of a higher spin theory. A model of second order in derivatives which describes a massive spin-3 particle is given by Singh and Hagen [7]. Here, we revisit the particle content of this model in  $D$  dimensions by suggesting a complete set of spin-projection operators; our results in some sense generalize the discussion carried out by [8] and are in agreement with those results obtained by [9]. We also provide a unitarity analysis of such model by adding source terms in order to verify the sign of the imaginary part of the residue of the transition amplitude saturated in the sources. For the massless case of the SH theory, we have obtained the canonical Hamiltonian as well as the constraints of the theory in  $D$  dimensions. By checking that they are first-class constraints, we demonstrate that the system describes the correct number of degrees of freedom. Finally, by using the constraints as strong equalities, we provide the reduced Hamiltonian in terms of spin-projection operators demonstrating that the model carries only spin-3 particles and that it is positive definite.

## 2. Rank-3 Spin-Projection Operators

In the SH model, the spin-3 field is a totally symmetric field  $h_{\mu\nu\lambda}$  with the trace given by  $h_\lambda = \eta^{\mu\nu} h_{\mu\nu\lambda}$ . Along with this

work, we have used the mostly plus metric  $(-, +, +, \dots)$ . At least in  $D = 3 + 1$ , we should expect six projection operators once the field  $h_{\mu\nu\lambda}$  belongs to the representation of the Lorentz group given by  $(1/2; 1/2) \otimes (1/2; 1/2) \otimes (1/2; 1/2) = (3/2; 3/2) \oplus (1/2; 1/2)$ . They correspond to the unique spin-3 sector given by the symmetric, transverse, and traceless part of  $h_{\mu\nu\lambda}$ , one spin-2 sector given by the divergence  $\partial^\mu h_{\mu\nu\lambda}$ , two spin-1 sectors contained at the double divergence  $\partial^\mu \partial^\nu h_{\mu\nu\lambda}$  and trace  $\eta^{\mu\nu} h_{\mu\nu\lambda}$ , and two spin-0 sectors given by the triple divergence  $\partial^\mu \partial^\nu \partial^\lambda h_{\mu\nu\lambda}$  and the divergence of the trace  $\partial^\mu h_\mu$ . Such projectors in this specific dimension are given for example in [8]. (A slightly different basis can also be found, for example, in [10], and a set of semi-projectors operators was explored by [11]. Both bases are not convenient to our purposes; besides, they are also in  $D = 3 + 1$ . A recent development has also been achieved for the rank- $s$  case in  $D$  dimensions at [12]). Aiming the construction of them in  $D$  dimensions, we notice that the trace of  $\theta_{\mu\nu} = \eta_{\mu\nu} - \omega_{\mu\nu}$  and  $\omega_{\mu\nu} = \partial_\mu \partial_\nu / \square$  are, respectively,  $D - 1$  and 1. Then one can generalize that the results to  $D$  dimensions are as follows:

$$\begin{aligned}
\left(P_{11}^{(3)}\right)_{\alpha\beta\gamma}^{\mu\nu\rho} &= \theta_{(\alpha}^{\mu} \theta_{\beta}^{\nu} \theta_{\gamma)}^{\rho} - \left(P_{11}^{(1)}\right)_{\alpha\beta\gamma}^{\mu\nu\rho}, \\
\left(P_{11}^{(2)}\right)_{\alpha\beta\gamma}^{\mu\nu\rho} &= 3\theta_{(\alpha}^{\mu} \theta_{\beta}^{\nu} \omega_{\gamma)}^{\rho} - \left(P_{11}^{(0)}\right)_{\alpha\beta\gamma}^{\mu\nu\rho}, \\
\left(P_{11}^{(1)}\right)_{\alpha\beta\gamma}^{\mu\nu\rho} &= \frac{3}{(D+1)} \theta^{(\mu\nu} \theta_{(\alpha\beta} \theta_{\gamma)}^{\rho)}, \\
\left(P_{22}^{(1)}\right)_{\alpha\beta\gamma}^{\mu\nu\rho} &= 3\theta_{(\alpha}^{\mu} \omega_{\beta}^{\nu} \omega_{\gamma)}^{\rho}, \\
\left(P_{11}^{(0)}\right)_{\alpha\beta\gamma}^{\mu\nu\rho} &= \frac{3}{(D-1)} \theta^{(\mu\nu} \theta_{(\alpha\beta} \omega_{\gamma)}^{\rho)}, \\
\left(P_{22}^{(0)}\right)_{\alpha\beta\gamma}^{\mu\nu\rho} &= \omega_{\alpha\beta} \omega^{\mu\nu} \omega_{\gamma}^{\rho}.
\end{aligned} \tag{1}$$

Notice that, here, the parenthesis means normalized symmetrization, taking, for example,

$$\begin{aligned}
\theta_{(\alpha}^{\mu} \theta_{\beta}^{\nu} \theta_{\gamma)}^{\rho} &= \frac{1}{6} \left( \theta_{\alpha}^{\mu} \theta_{\beta}^{\nu} \theta_{\gamma}^{\rho} + \theta_{\alpha}^{\rho} \theta_{\beta}^{\nu} \theta_{\gamma}^{\mu} + \theta_{\alpha}^{\nu} \theta_{\beta}^{\mu} \theta_{\gamma}^{\rho} \right. \\
&\quad \left. + \theta_{\alpha}^{\rho} \theta_{\beta}^{\mu} \theta_{\gamma}^{\nu} + \theta_{\alpha}^{\nu} \theta_{\beta}^{\rho} \theta_{\gamma}^{\mu} + \theta_{\alpha}^{\mu} \theta_{\beta}^{\rho} \theta_{\gamma}^{\nu} \right).
\end{aligned} \tag{2}$$

As a requirement, the basis must be orthonormal and the projector idempotents:

$$P_{ij}^{(s)} P_{kl}^{(r)} = \delta^{sr} \delta_{jk} P_{il}^{(s)}. \tag{3}$$

In our notation, the superscripts  $(r)$  and  $(s)$  denote the spin subspace, while subscripts  $i, j, k$ , and  $l$  work to distinguish between projectors and transition operators. Once, for example,  $i = j$  or  $k = l$ , we have a projector, while if  $i \neq j$  or  $k \neq l$ , we have a transition operator. In addition, the subscripts work in order to count the number of projectors of a given spin subspace, for example, in the subspace of spin 0, we have

two projectors represented by the combinations  $i = 1, j = 1$  and  $i = 2, j = 2$ . The set of projectors obeys the following mathematical identity:

$$\sum_{i,s} P_{ii}^{(s)} = \mathbb{1}, \tag{4}$$

where  $\mathbb{1}$  stands for the symmetric rank-3 identity operator, i.e.,

$$\mathbb{1}_{\alpha\beta\gamma}^{\mu\nu\rho} = \delta_{(\alpha}^{\mu} \delta_{\beta}^{\nu} \delta_{\gamma)}^{\rho)}. \tag{5}$$

Finally, the transition operators  $P_{ij}^{(s)}$  are given by

$$\begin{aligned}
\left(P_{12}^{(1)}\right)_{\alpha\beta\gamma}^{\mu\nu\rho} &= \frac{3}{\sqrt{(D+1)}} \theta_{(\alpha\beta} \theta_{\gamma)}^{\rho} \omega^{\mu\nu}, \\
\left(P_{21}^{(1)}\right)_{\alpha\beta\gamma}^{\mu\nu\rho} &= \frac{3}{\sqrt{(D+1)}} \theta^{(\mu\nu} \theta_{\gamma)}^{\rho} \omega_{\alpha\beta}, \\
\left(P_{12}^{(0)}\right)_{\alpha\beta\gamma}^{\mu\nu\rho} &= \frac{3}{\sqrt{3(D-1)}} \theta_{(\alpha\beta} \omega_{\gamma)}^{\rho} \omega^{\mu\nu}, \\
\left(P_{21}^{(0)}\right)_{\alpha\beta\gamma}^{\mu\nu\rho} &= \frac{3}{\sqrt{3(D-1)}} \theta^{(\mu\nu} \omega_{(\alpha\beta} \omega_{\gamma)}^{\rho)}.
\end{aligned} \tag{6}$$

We have to say that the transition operators also satisfy the algebra given by (3). They are not necessary to complete the identity in (4); however, we do need them in order to expand the sandwiched operator between two rank-3 fields in a bilinear Lagrangian.

### 3. On the Coefficients of the Spin-3 Singh-Hagen Theory

As showed in [13, 14], the description of a massive spin-3 particle in terms of a totally symmetric field with a second-order Lagrangian requires the introduction of an auxiliary field. The simplest way to introduce it is by associating the totally symmetric field with a scalar field. In [15], one of us in collaboration has observed that even for higher derivative descriptions of doublets of spin-3, the addition of such auxiliaries is necessary. Recently, however, [16] have verified that they are not necessary for a higher derivative self-dual description of a massive spin-3 singlet in  $D = 2 + 1$  dimensions.

In the following lines, we use the operators to revisit, as an example, the particle content of the SH model coupled in the simplest way to a scalar auxiliary field  $W$ . Let us suppose all the coefficients are undetermined and given by  $a, b, \dots, t$ , then we have

$$\begin{aligned}
\mathcal{L} &= a h_{\mu\nu\rho} \square h^{\mu\nu\rho} + b h_{\mu\nu\rho} \partial^\mu \partial_\alpha h^{\alpha\nu\rho} + c h_\nu \partial^\nu \partial_\mu h^\mu + d h_\nu \square h^\nu \\
&\quad + e h_\rho \partial_\mu \partial_\nu h^{\mu\nu\rho} + f m^2 h_{\mu\nu\rho} h^{\mu\nu\rho} + g m^2 h_\mu h^\mu + i W \square W \\
&\quad + j m^2 W^2 + t m h_\mu \partial^\mu W.
\end{aligned} \tag{7}$$

In the appendix, we give explicit expressions of the bilinear form of each term. By collecting all of them, the Lagrangian density can be written as

$$\mathcal{L} = h_{\mu\nu\rho} \Theta_{\kappa\lambda\sigma}^{\mu\nu\rho} h^{\kappa\lambda\sigma} + W\Phi W + h_{\mu\nu\rho} T^{\mu\nu\rho} W, \quad (8)$$

where

$$\begin{aligned} \Theta_{\kappa\lambda\sigma}^{\mu\nu\rho} = & (a\Box + fm^2) (P_{11}^{(3)})_{\kappa\lambda\sigma}^{\mu\nu\rho} + \left[ \left( a + \frac{b}{3} \right) \Box + fm^2 \right] (P_{11}^{(2)})_{\kappa\lambda\sigma}^{\mu\nu\rho} \\ & + \left\{ \left[ a + \frac{(D+1)}{3} d \right] \Box + \left[ f + \frac{(D+1)}{3} g \right] m^2 \right\} (P_{11}^{(1)})_{\kappa\lambda\sigma}^{\mu\nu\rho} \\ & + \left[ \left( a + \frac{2b+d+e}{3} \right) \Box + \left( f + \frac{g}{3} \right) m^2 \right] (P_{22}^{(1)})_{\kappa\lambda\sigma}^{\mu\nu\rho} \\ & + \left\{ \left[ a + \frac{b+(D-1)(c+d)}{3} \right] \Box + \left[ f + \frac{(D-1)}{3} g \right] m^2 \right\} \\ & \cdot (P_{11}^{(0)})_{\kappa\lambda\sigma}^{\mu\nu\rho} + [(a+b+c+d+e)\Box + (f+g)m^2] (P_{22}^{(0)})_{\kappa\lambda\sigma}^{\mu\nu\rho} \\ & + \frac{\sqrt{D+1}}{3} \left[ \left( d + \frac{e}{2} \right) \Box + gm^2 \right] (P_{12}^{(1)} + P_{21}^{(1)})_{\kappa\lambda\sigma}^{\mu\nu\rho} \\ & + \frac{\sqrt{3(D-1)}}{3} \left[ \left( c + d + \frac{e}{2} \right) \Box + gm^2 \right] (P_{12}^{(0)} + P_{21}^{(0)})_{\kappa\lambda\sigma}^{\mu\nu\rho}, \end{aligned}$$

$$\Phi = (i\Box + jm^2),$$

$$T^{\mu\nu\rho} = \frac{tm}{3} (\eta^{\mu\nu} \partial^\rho + \eta^{\nu\rho} \partial^\mu + \eta^{\rho\mu} \partial^\nu). \quad (9)$$

Integrating over the scalar field  $W$ , we then have the nonlocal Lagrangian given by

$$\mathcal{L} = h_{\mu\nu\rho} G_{\alpha\beta\gamma}^{\mu\nu\rho} h^{\alpha\beta\gamma}, \quad (10)$$

where the operator  $G$  is written as

$$G_{\kappa\lambda\sigma}^{\mu\nu\rho} \equiv \Theta_{\kappa\lambda\sigma}^{\mu\nu\rho} + \frac{1}{4\Phi} T^{\mu\nu\rho} T_{\kappa\lambda\sigma}. \quad (11)$$

In order to determine the coefficients  $a, b, \dots, t$ , we now take the equations of motion with respect to the symmetric field  $h^{\mu\nu\lambda}$ , and then to select the spin-3, spin-2, spin-1, and spin-0 sectors, we apply the spin-projection operators on these equations, which give us

$$(P_{11}^{(3)})_{\mu\nu\rho}^{\alpha\beta\gamma} (Gh)^{\mu\nu\rho} = (a\Box + fm^2) (P_{11}^{(3)} h)^{\alpha\beta\gamma}, \quad (12)$$

$$(P_{11}^{(2)})_{\mu\nu\rho}^{\alpha\beta\gamma} (Gh)^{\mu\nu\rho} = \left[ \left( a + \frac{b}{3} \right) \Box + fm^2 \right] (P_{11}^{(2)} h)^{\alpha\beta\gamma}, \quad (13)$$

$$\begin{aligned} (P_{11}^{(1)})_{\mu\nu\rho}^{\alpha\beta\gamma} (Gh)^{\mu\nu\rho} = & \left\{ \left[ a + \frac{(D+1)}{3} d \right] \Box + \left[ f + \frac{(D+1)}{3} g \right] m^2 \right\} \\ & \cdot (P_{11}^{(1)} h)^{\alpha\beta\gamma} + \frac{\sqrt{D+1}}{3} \left[ \left( d + \frac{e}{2} \right) \Box + gm^2 \right] \\ & \cdot (P_{12}^{(1)} P_{22}^{(1)} h)^{\alpha\beta\gamma}, \end{aligned} \quad (14)$$

$$\begin{aligned} (P_{22}^{(1)})_{\mu\nu\rho}^{\alpha\beta\gamma} (Gh)^{\mu\nu\rho} = & \frac{\sqrt{D+1}}{3} \left[ \left( d + \frac{e}{2} \right) \Box + gm^2 \right] (P_{21}^{(1)} P_{11}^{(1)} h)^{\alpha\beta\gamma} \\ & + \left[ \left( a + \frac{2b+d+e}{3} \right) \Box + \left( f + \frac{g}{3} \right) m^2 \right] \\ & \cdot (P_{22}^{(1)} h)^{\alpha\beta\gamma}, \end{aligned} \quad (15)$$

$$\begin{aligned} (P_{11}^{(0)})_{\mu\nu\rho}^{\alpha\beta\gamma} (Gh)^{\mu\nu\rho} = & \left\{ \left[ a + \frac{b}{3} + \frac{(D-1)(c+d)}{3} \right] \Box \right. \\ & + \left. \left[ f + \frac{(D-1)}{3} g \right] m^2 \right\} (P_{11}^{(0)} h)^{\alpha\beta\gamma} \\ & + \frac{\sqrt{3(D-1)}}{3} \left[ \left( c + d + \frac{e}{2} \right) \Box + gm^2 \right] \\ & \cdot (P_{12}^{(0)} P_{22}^{(0)} h)^{\alpha\beta\gamma} + \frac{t^2 m^2 \Box}{12(i\Box + jm^2)} \\ & \cdot \left[ (D-1) (P_{11}^{(0)} h)^{\alpha\beta\gamma} \right. \\ & \left. + \sqrt{3(D-1)} (P_{12}^{(0)} P_{22}^{(0)} h)^{\alpha\beta\gamma} \right], \end{aligned} \quad (16)$$

$$\begin{aligned} (P_{22}^{(0)})_{\mu\nu\rho}^{\alpha\beta\gamma} (Gh)^{\mu\nu\rho} = & \frac{\sqrt{3(D-1)}}{3} \left[ \left( c + d + \frac{e}{2} \right) \Box + gm^2 \right] \\ & \cdot (P_{21}^{(0)} P_{11}^{(0)} h)^{\alpha\beta\gamma} + [(a+b+c+d+e)\Box \\ & + (f+g)m^2] (P_{22}^{(0)} h)^{\alpha\beta\gamma} + \frac{t^2 m^2 \Box}{12(i\Box + jm^2)} \\ & \cdot \left[ \sqrt{3(D-1)} (P_{21}^{(0)} P_{11}^{(0)} h)^{\alpha\beta\gamma} + 3 (P_{22}^{(0)} h)^{\alpha\beta\gamma} \right]. \end{aligned} \quad (17)$$

Once we want to have only the spin-3 propagation, we need to handle the equations from (12) to (17) in order to have a Klein-Gordon equation in the spin-3 sector and to kill all the subsidiary conditions propagating lower spins, which can be done by setting to zero the coefficients multiplying the d'Alembertian in such sectors. After manipulating with the system of equations given above, one can then find the coefficients and their relations, which after substituting back in the Lagrangian density give us

$$\begin{aligned} \mathcal{L} = & a h_{\mu\nu\rho} \Box h^{\mu\nu\rho} - 3a h_{\mu\nu\rho} \partial^\mu \partial_\alpha h^{\alpha\nu\rho} - \frac{3a}{2} h_{\nu} \partial^\nu \partial_\mu h^\mu \\ & - 3a h_{\nu} \Box h^\nu + 6a h_\rho \partial_\mu \partial_\nu h^{\mu\nu\rho} - am^2 h_{\mu\nu\rho} h^{\mu\nu\rho} \\ & + 3am^2 h_\mu h^\mu - \frac{1}{3a} \frac{(D-1)}{(D-2)} t^2 W \Box W \\ & + \frac{1}{2a} \left( \frac{D}{D-2} \right)^2 t^2 m^2 W^2 + tm h_\mu \partial^\mu W. \end{aligned} \quad (18)$$

We end up with two undetermined coefficients  $a$  and  $t$ ; however, notice that if we redefine  $h \rightarrow h/\sqrt{2a}$  and  $W \rightarrow \sqrt{2a}W/t$ , they are completely eliminated of the action. Besides, we notice that the dimensional dependence in the coefficients of  $W\Box W$  and  $W^2$  is exactly the same

with what Aragone et al. have obtained[9]. This is precisely the SH model in  $D$  dimensions, and in the next section, we are going to analyze its unitarity. The presence of auxiliary fields in higher spin theories is always the reason for difficulties when analyzing the equations of motion. We have noticed that if one takes the Lagrangian given by (7) and eliminate the scalar field ad hoc, we can then perform the analysis of the equations of motion by projecting the spin sectors and conclude that we have a Klein-Gordon equation to the spin-3 mode if and only if  $D=2$ . In other words, in this specific dimension, we do not need the presence of auxiliary fields. This is an expected result once there is no reason to think about spins in such a dimension.

#### 4. Unitarity of the Spin-3 Singh-Hagen Model

Here, we start by supposing that one can integrate over the auxiliary scalars obtaining a nonlocal Lagrangian which however can be put in a bilinear form. Then the sandwiched operator can be expanded in terms of the orthonormal basis introduced before. As a warm-up exercise, we could take a general bilinear Lagrangian given by

$$\mathcal{L} = h_{\mu\nu\rho} G_{\alpha\beta\gamma}^{\mu\nu\rho} h^{\alpha\beta\gamma}, \quad (19)$$

where  $G_{\alpha\beta\gamma}^{\mu\nu\rho}$  is an operator that can be written in terms of  $P_{ij}^{(s)}$ . From now on, we have suppressed the indices for the sake of simplicity in all the results.

$$G = AP_{11}^{(3)} + BP_{11}^{(2)} + CP_{11}^{(1)} + DP_{22}^{(1)} + E(P_{12}^{(1)} + P_{21}^{(1)}) + FP_{11}^{(0)} + HP_{22}^{(0)} + I(P_{12}^{(0)} + P_{21}^{(0)}), \quad (20)$$

and suppose the inverse of  $G$  given by

$$G^{-1} = JP_{11}^{(3)} + KP_{11}^{(2)} + LP_{11}^{(1)} + MP_{22}^{(1)} + N(P_{12}^{(1)} + P_{21}^{(1)}) + QP_{11}^{(0)} + RP_{22}^{(0)} + S(P_{12}^{(0)} + P_{21}^{(0)}). \quad (21)$$

The coefficients  $A, B, \dots, S$  are completely arbitrary, but once we know that  $GG^{-1} = 1$ , we can relate them through the general results below:

$$\begin{aligned} J &= \frac{1}{A}; K = \frac{1}{B}, \\ L &= \frac{D}{DC - E^2}; M = \frac{C}{DC - E^2}; N = -\frac{E}{DC - E^2}, \\ Q &= \frac{H}{HF - I^2}; R = \frac{F}{HF - I^2}; S = -\frac{I}{HF - I^2}. \end{aligned} \quad (22)$$

Such results are useful for any spin-3 model without parity breaking. This is precisely the case of the Lagrangian density we have found at (18) after field redefinitions. One can verify that the operator  $G$  in this case is given by

$$\begin{aligned} G &= \frac{(\square - m^2)}{2} (P_{11}^{(3)}) - \frac{m^2}{2} (P_{11}^{(2)}) - \frac{D}{2} (\square - m^2) (P_{11}^{(1)}) \\ &\quad - \frac{1}{2} \left[ \frac{3(D-1)\square}{2} - (D-2)m^2 \right] (P_{11}^{(0)}) \\ &\quad - \frac{1}{2} \left[ \frac{\square}{2} - 2m^2 \right] (P_{22}^{(0)}) + \frac{1}{2} \sqrt{D+1} m^2 (P_{12}^{(1)} + P_{21}^{(1)}) \\ &\quad - \frac{1}{2} \sqrt{3(D-1)} \left( \frac{\square}{2} - m^2 \right) (P_{12}^{(0)} + P_{21}^{(0)}) \\ &\quad + \frac{m^2 \square}{12(i\square + jm^2)} \left[ (D-1) (P_{11}^{(0)}) + 3 (P_{22}^{(0)}) \right] \\ &\quad + \frac{m^2 \sqrt{3(D-1)} \square}{12(i\square + jm^2)} (P_{12}^{(0)} + P_{21}^{(0)}) \end{aligned} \quad (23)$$

where we have used  $i = -2(D-1)/3(D-2)$  and  $j = D^2/(D-2)^2$ . With the help of the general expressions obtained before, after inverting the operator  $G$ , we have the following propagator:

$$\begin{aligned} (G^{-1}) &= \frac{2}{(\square - m^2)} (P_{11}^{(3)}) - \frac{2}{m^2} (P_{11}^{(2)}) + \frac{2D(\square - m^2)}{(D+1)m^4} (P_{22}^{(1)}) + \frac{2\sqrt{D+1}}{(D+1)m^2} (P_{12}^{(1)} + P_{21}^{(1)}) \\ &\quad - \frac{2}{3} \left[ \frac{(D-1)(D-2)\square^2 - 2(D-1)(2D-1)\square m^2 + 6D^2 m^4}{D^2(D+1)m^6} \right] (P_{11}^{(0)}) \\ &\quad - 2 \left[ \frac{(D-1)^2(D-2)\square^2 - 2(D-1)(D^2 - D + 1)\square m^2 + D^2(D-2)m^4}{D^2(D+1)m^6} \right] (P_{22}^{(0)}) \\ &\quad + \frac{2}{3} \sqrt{3(D-1)} \left[ \frac{(D-1)(D-2)\square^2 - (3D^2 - 4D + 2)\square m^2 + 3D^2 m^4}{D^2(D+1)m^6} \right] (P_{12}^{(0)} + P_{21}^{(0)}). \end{aligned} \quad (24)$$

Notice that there is a massive pole in the spin-3 sector and we have no dynamics in the lower spin sectors. In order to check that this spin-3 particle is, in fact, a physical particle, let us analyze the sign of the imaginary part of the residue of the transition amplitude saturated in the sources. After taking the Fourier transformation, in the momentum space, it is given by

$$\mathcal{A}(k) = -\frac{i}{2} F_{\mu\nu\rho}^*(k) (G^{-1}(k))_{\alpha\beta\gamma}^{\mu\nu\rho} F^{\alpha\beta\gamma}(k), \quad (25)$$

where  $F_{\mu\nu\rho}(k)$  is the source in such space. We have a physical particle if the following condition on the transition amplitude is satisfied:

$$\text{Im} [\text{Res}(\mathcal{A}_2(k))|_{k^2=-m^2}] > 0. \quad (26)$$

Otherwise, we would have a ghost at the spectrum. In our case, we have

$$\text{Im} [\text{Res}(\mathcal{A}(k))|_{k^2=-m^2}] = \lim_{k^2 \rightarrow -m^2} (k^2 + m^2) \mathcal{A}(k) = \tilde{F}_{\mu\nu\rho}^* \tilde{F}^{\mu\nu\rho}, \quad (27)$$

where  $\tilde{F}_{\mu\nu\rho} = (P_{11}^{(3)})_{\mu\nu\rho}^{\alpha\beta\gamma} \tilde{F}_{\alpha\beta\gamma}$ . As the spin-3 projection operator  $(P_{11}^{(3)})$  is totally symmetric with respect to the indices  $\mu\nu\rho$  and  $\alpha\beta\gamma$ , traceless, and transverse, the source term must have the same properties, i.e.,

$$\begin{aligned} \eta^{\mu\nu} \tilde{F}_{\mu\nu\rho} &= 0, \\ k^\mu \tilde{F}_{\mu\nu\rho} &= 0. \end{aligned} \quad (28)$$

Once we have only one massive pole at the spin-3 sector, we choose the frame given by  $k^\mu = (m, 0, \dots, 0_{(D-1)})$ . Then, we have

$$k^\mu \tilde{F}_{\mu\nu\rho} = m \tilde{F}_{0\nu\rho} = 0 \Rightarrow \tilde{F}_{0\nu\rho} = 0, \quad (29)$$

which leaves us only with the spatial contributions, given by

$$\text{Im} [\text{Res}(\mathcal{A})|_{k^2=-m^2}] = \tilde{F}_{ijl}^* \tilde{F}^{ijl} > 0; \quad i, j, k = 1, \dots, D-1. \quad (30)$$

From this spectral analysis, we can verify that the SH model is free of ghosts, propagating only a spin-3 massive particle in  $D$  dimensions. It is also interesting to notice that once the propagator of the SH theory (24) has a unique pole in the spin-3 sector, which goes with  $1/(k^2 + m^2)$  in the momentum space, the potential in the nonrelativistic limit (at least in  $D=3+1$ ) between two sources exchanging spin-3 particles is a Yukawa potential  $V(r) \sim e^{-mr}/r$ , as also happens with the lower spin cases, namely, Proca and *FP* theories. The massless limit of higher spin theories however may have discontinuities, as observed by van Dam and Veltman for spin-2 theories [19] and by Berends and Reisen [8] for the case of spin-3 theories.

## 5. Hamiltonian Positivity of the Massless Case

The massless SH theory is given by

$$\begin{aligned} L_{SH}^{m=0} &= -\frac{1}{2} (\partial_\alpha h_{\mu\nu\rho})^2 + \frac{3}{2} (\partial^\mu h_{\mu\nu\rho})^2 + \frac{3}{2} (\partial_\alpha h_\nu)^2 + \frac{3}{4} (\partial^\nu h_\nu)^2 \\ &\quad - 3 \partial_\nu h_\rho \partial_\mu h^{\mu\nu\rho}. \end{aligned} \quad (31)$$

In order to study the Hamiltonian positivity of this system, we break manifest Lorentz covariance identifying terms with two time derivatives, which give us

$$\begin{aligned} L_{SH}^{m=0} &= \frac{1}{2} (\dot{h}_{ijk})^2 - \frac{3}{2} (\dot{\bar{h}}_i)^2 + \frac{1}{4} \dot{\lambda}^2 + \frac{3}{2} \dot{\lambda} \partial_i h_{00i} - 3 \partial_k h_{0ij} \dot{h}_{ijk} \\ &\quad + \frac{3}{2} \partial_i \lambda \dot{\bar{h}}_i + 3 \partial_i \bar{h}_0 \dot{\bar{h}}_i + 6 \partial_i h_{0ij} \dot{\bar{h}}_j + \mathcal{V}, \end{aligned} \quad (32)$$

where we have used the following notation  $\bar{h}_i = h_{ijj}$ ,  $\bar{h}_0 = h_{0kk}$ , and  $\lambda = h_{000} - 3\bar{h}_0$ , which is similar to the one adopted by [20]. Notice also that all the terms without time derivatives are grouped on the symbol  $\mathcal{V}$  which is explicitly given by

$$\begin{aligned} \mathcal{V} &= -(\partial_l \lambda)^2 - 3 \partial_l \lambda \partial_l \bar{h}_0 - \frac{3}{2} (\partial_l \bar{h}_0)^2 + \frac{3}{2} (\partial_l h_{0ij})^2 \\ &\quad - \frac{1}{2} (\partial_l h_{ijk})^2 + \frac{9}{4} (\partial_i h_{00i})^2 - 3 (\partial_i h_{0ij})^2 + \frac{3}{2} (\partial_i h_{ijk})^2 \\ &\quad + \frac{3}{2} (\partial_i \bar{h}_i)^2 + \frac{3}{4} (\partial_i \bar{h}_i)^2 - 3 \partial_i h_{00i} \partial_i \bar{h}_i - \frac{3}{2} \partial_i h_{00i} \partial_i \bar{h}_i \\ &\quad + 3 \partial_j h_{k00} \partial_i h_{ijk} - 3 \partial_j \bar{h}_k \partial_i h_{ijk} - 3 \partial_i h_{0ij} \partial_j \lambda - 6 \partial_j \bar{h}_0 \partial_i h_{0ij}. \end{aligned} \quad (33)$$

Calculating the conjugate momenta from (32), we have the following primary constraints  $\varphi^i = \pi^{00i} \approx 0$  and  $\varphi^{ij} = \pi^{0ij} \approx 0$ . The consistency checking *a la* the Dirac-Bergman procedure of such constraints gives us two additional secondary constraints  $\chi^i$  and  $\chi^{ij}$  which are given by

$$\begin{aligned} \chi^i &= 3 \left( \partial^i \pi - \nabla^2 \bar{h}^i + \partial_j \partial_k h^{ijk} - \frac{1}{2} \partial^i \partial_j \bar{h}^j \right), \\ \chi^{ij} &= -3 \left( \partial_k \pi^{kij} + \frac{1}{2} \delta^{ij} \nabla^2 \lambda \right). \end{aligned} \quad (34)$$

One can verify that all the constraints are first class, which is related to the fact that the massless theory is gauge invariant under traceless reparametrizations, i.e.,  $\delta h_{\mu\nu\lambda} = \partial_{(\mu} \tilde{\xi}_{\nu\lambda)}$  where  $\tilde{\xi}_{\nu\lambda} = \tilde{\xi}_{\lambda\nu}$  and  $\eta^{\nu\lambda} \tilde{\xi}_{\nu\lambda} = 0$ . Then in the Lagrangian level, once we have two vectorial and two tensorial first-class constraints, by considering the number of independent components of the totally symmetric rank-3 tensor, one has  $D(D+1)(D+2)/6 - 2D(D-1)/2 - 2(D-1) = (D-2)(D-3)(D+2)/6$  degrees of freedom which correspond to the two helicities  $+3$  and  $-3$  in  $D=4$ . With these results in hand, we then take the Legendre transformation, in



order to obtain the canonical Hamiltonian, which can be written as

$$\begin{aligned} \mathcal{H} = & \pi^2 + \frac{1}{2} (\pi^{ijk})^2 - \frac{3}{2D} (\pi^k)^2 + \frac{3}{2D} \pi^k \partial_k \lambda \\ & + \left( \frac{5D-3}{8D} \right) (\partial_k \lambda)^2 + \frac{1}{2} (\partial_l h_{ijk})^2 - \frac{3}{2} (\partial_l h_{ijk})^2 \\ & + 3 \partial_l h_{ijk} \partial_j \bar{h}_k - \frac{3}{2} (\partial_l \bar{h}_i)^2 - \frac{3}{4} (\partial_i \bar{h}_i)^2 + h_{00i} \chi^i + h_{0ij} \chi^{ij}. \end{aligned} \quad (35)$$

Here, the momentum  $\pi$  is canonically conjugated to the combination  $\lambda$ . Following the same approach used by [21], one can now determine the partially reduced Hamiltonian, which is done by using the secondary constraints as strong equalities, which give us

$$\begin{aligned} \mathcal{H}^{(pr)} = & \pi_{ijk} \left[ \frac{1}{2} (P_{11}^{(3)} + P_{11}^{(2)}) + \frac{(D-1)}{2D} P_{22}^{(1)} - \frac{\sqrt{D}}{2D} (P_{12}^{(1)} + P_{21}^{(1)}) \right. \\ & + \frac{(5D-7)}{6(D-1)^2} P_{11}^{(0)} + \frac{(D-2)(D-3)}{2(D-1)^2} P_{22}^{(0)} \\ & \left. - \frac{(3D-5)}{6(D-1)^2} \sqrt{3(D-2)} (P_{12}^{(0)} + P_{21}^{(0)}) \right]_{lmn}^{ijk} \pi^{lmn} \\ & + h_{ijk} \left[ \nabla^2 \left( -\frac{1}{2} P_{11}^{(3)} + \frac{(D-1)}{2} P_{11}^{(1)} \right) \right]_{lmn}^{ijk} h^{lmn}, \end{aligned} \quad (36)$$

where we have written all the bilinear contractions of rank-3 tensors in terms of the spin-projection and the transition operators (Notice that we need to redefine our projectors, changing  $D \rightarrow D-1$ . Besides, the transverse and longitudinal operator are, respectively, given by  $\theta_{ij} = \delta_{ij} - \partial_i \partial_j / \nabla^2$  and  $\omega_{ij} = \partial_i \partial_j / \nabla^2$ ). By using the secondary constraints, one can then get rid of the terms proportional to  $P_{11}^{(1)}$ ,  $P_{11}^{(2)}$ ,  $P_{22}^{(1)}$ , and  $(P_{12}^{(1)} + P_{21}^{(1)})$  separately. Notice also that the spin-0 sector can be combined in order to vanish. This leads us with the reduced Hamiltonian given only in terms of the spin-3 projection operators, as expected for the massless spin  $\pm 3$  particles described by the kinetic part of the SH theory.

$$\mathcal{H}^{(r)} = \frac{1}{2} \pi_{ijk} \left( P_{11}^{(3)} \right)_{lmn}^{ijk} \pi^{lmn} - \frac{1}{2} h_{ijk} \left( \nabla^2 P_{11}^{(3)} \right)_{lmn}^{ijk} h^{lmn} \geq 0. \quad (37)$$

Once any projection operator sandwiched among two equal tensors has positive definite bilinear, and since  $-\nabla^2$  has only positive eigenvalues, one can conclude that the reduced Hamiltonian is positive definite implying in its classical stability.

## 6. Conclusion

Here, we provide a generalized set of spin-projection and transition operators for symmetric rank-3 tensors in  $D$  dimensions. By setting  $D=3+1$ , we can recover the results of [8], and in this sense, we have a generalization of those projectors. In a work in progress, we are constructing a set of spin-projection operators to the case of nonsymmetric fields in  $D$  dimensions, and in addition, we have considered the case of models with parity breaking in  $D=2+1$ , as it is the case of the self-dual models suggested by [16, 17].

The coefficients of the SH model are determined with the help of the spin-projection operators, by projecting the equations of motion in the subspace of spin-3 and lower. We have noticed that we end up with two arbitrary coefficients  $a$  and  $t$ . But both of them may be eliminated through field redefinitions. In a work in progress, we have been analyzing the particle content of the higher derivative models obtained at [15], but an additional difficulty is in the game now; once the models are gauge invariants, we need to construct gauge-fixing terms, in order to obtain the propagator. Besides, as the models have higher derivatives, double poles should appear. It is interesting to say that in  $D=2+1$ , such higher derivative descriptions have lots of similarities with the new massive gravity model for spin-2 particles [18].

Finally, we have given a detailed obtainment of the propagator of the SH model by means of a general expression for bilinear Lagrangians in terms of totally symmetric rank-3 tensors. From this analysis, we can verify that there is only a spin-3 particle in the spectrum of the theory. By adding a source term, we have analyzed the sign of the imaginary part of the residue of the transition amplitude saturated, and as it was expected, the propagating mode is physical. In a way quite similar to the analysis performed by [21], we have also obtained the Hamiltonian as well as the constraints of the massless SH theory in  $D$  dimensions. Thanks to the projection and transition operators, we have then demonstrated that the system is positive definite, describing correctly the number of degrees as well as the spin.

## Appendix

Here, we provide to the reader a detailed set of equations step by step where we write the terms in their bilinear forms in terms of the spin-projection operators:

$$\begin{aligned} h_{\mu\nu\rho} \square h^{\mu\nu\rho} = & h_{\mu\nu\rho} \left[ \square \left( P_{11}^{(3)} + P_{11}^{(2)} + P_{11}^{(1)} \right. \right. \\ & \left. \left. + P_{22}^{(1)} + P_{11}^{(0)} + P_{22}^{(0)} \right)^{\mu\nu\rho} \right]_{\kappa\lambda\sigma} h^{\kappa\lambda\sigma} \end{aligned}$$

$$\begin{aligned} h_{\mu\nu\rho} \partial^\mu \partial_\alpha h^{\alpha\nu\rho} = & h_{\mu\nu\rho} \left[ \frac{\square}{18} \left( 6P_{11}^{(2)} + 12P_{22}^{(1)} \right. \right. \\ & \left. \left. + 6P_{11}^{(0)} + 18P_{22}^{(0)} \right)^{\mu\nu\rho} \right]_{\kappa\lambda\sigma} h^{\kappa\lambda\sigma} \end{aligned}$$

$$\begin{aligned}
h_\nu \partial^\nu \partial_\mu h^\mu &= h_{\mu\nu\rho} \left[ \frac{\square}{9} \left( 3(D-1)P_{11}^{(0)} + 9P_{22}^{(0)} \right. \right. \\
&\quad \left. \left. + 3\sqrt{3(D-1)} \left( P_{12}^{(0)} + P_{21}^{(0)} \right) \right)^{\mu\nu\rho} \right] h^{\kappa\lambda\sigma} \\
h_\nu \square h^\nu &= h_{\mu\nu\rho} \left[ \frac{\square}{9} \left( 3(D+1)P_{11}^{(1)} + 3P_{22}^{(1)} + 3(D-1)P_{11}^{(0)} \right. \right. \\
&\quad \left. \left. + 9P_{22}^{(0)} \right)^{\mu\nu\rho} \right] h^{\kappa\lambda\sigma} + h_{\mu\nu\rho} \left[ \frac{\square}{9} \left( 3\sqrt{D+1} \left( P_{12}^{(1)} \right. \right. \right. \\
&\quad \left. \left. + P_{21}^{(1)} \right) + 3\sqrt{3(D-1)} \left( P_{12}^{(0)} + P_{21}^{(0)} \right) \right)^{\mu\nu\rho} \right] h^{\kappa\lambda\sigma} \\
h_\rho \partial_\mu \partial_\nu h^{\mu\nu\rho} &= h_{\mu\nu\rho} \left[ \frac{\square}{18} \left( 6P_{22}^{(1)} + 18P_{22}^{(0)} + 3\sqrt{D+1} \left( P_{12}^{(1)} \right. \right. \right. \\
&\quad \left. \left. + P_{21}^{(1)} \right) \right)^{\mu\nu\rho} \right] h^{\kappa\lambda\sigma} + h_{\mu\nu\rho} \left[ \frac{\square}{18} \left( 3\sqrt{3(D-1)} \left( P_{12}^{(0)} \right. \right. \right. \\
&\quad \left. \left. + P_{21}^{(0)} \right) \right)^{\mu\nu\rho} \right] h^{\kappa\lambda\sigma} \\
m^2 h_{\mu\nu\rho} h^{\mu\nu\rho} &= h_{\mu\nu\rho} \left[ m^2 \left( P_{11}^{(3)} + P_{11}^{(2)} + P_{11}^{(1)} \right. \right. \\
&\quad \left. \left. + P_{22}^{(1)} + P_{11}^{(0)} + P_{22}^{(0)} \right)^{\mu\nu\rho} \right] h^{\kappa\lambda\sigma} \\
m^2 h_\mu h^\mu &= h_{\mu\nu\rho} \left[ \frac{m^2}{9} \left( 3(D+1)P_{11}^{(1)} + 3P_{22}^{(1)} + 3(D-1)P_{11}^{(0)} \right. \right. \\
&\quad \left. \left. + 9P_{22}^{(0)} \right)^{\mu\nu\rho} \right] h^{\kappa\lambda\sigma} + h_{\mu\nu\rho} \left[ \frac{m^2}{9} \left( 3\sqrt{D+1} \left( P_{12}^{(1)} \right. \right. \right. \\
&\quad \left. \left. + P_{21}^{(1)} \right) + 3\sqrt{3(D-1)} \left( P_{12}^{(0)} + P_{21}^{(0)} \right) \right)^{\mu\nu\rho} \right] h^{\kappa\lambda\sigma}.
\end{aligned} \tag{A.1}$$

Collecting all these terms, one can write the first and most tedious term of the Lagrangian density given by (8).

## Data Availability

No data were used to support this study.

## Conflicts of Interest

The authors declare that they have no conflicts of interest.

## Acknowledgments

This study was financed in part by the Coordenação de Aperfeiçoamento de Pessoal de Nível Superior-Brasil (CAPES)-Finance Code 001. We acknowledge helpful discussions with Prof. Denis Dalmazi.

## References

- [1] P. Dirac, "Relativistic wave equations," *Proceedings of the Royal Society of London: Series A, Mathematical and Physical Sciences*, vol. 155, no. 886, pp. 447–459, 1936.
- [2] M. Fierz and W. Pauli, "On relativistic wave equations for particles of arbitrary spin in an electromagnetic field," *Proceedings of the Royal Society of London: Series A, Mathematical and Physical Sciences*, vol. 173, no. 953, pp. 211–232, 1939.
- [3] R. J. Rivers, "Lagrangian theory of neutral massive spin-2 fields," *Il Nuovo Cimento*, vol. 34, p. 387, 1964.
- [4] P. van Nieuwenhuizen, "On ghost-free tensor Lagrangians and linearized gravitation," *Nuclear Physics B*, vol. 60, pp. 478–492, 1973.
- [5] F. C. P. Nunes and G. O. Pires, "Extending the Barnes-Rivers operators to  $D=3$  topological gravity," *Physics Letters B*, vol. 301, no. 4, pp. 339–344, 1993.
- [6] A. Accioly, J. Helayël-Neto, B. Pereira-Dias, and C. Hernaski, "New class of spin projection operators for 3D models," *Physical Review D*, vol. 86, no. 10, article 105046, 2012.
- [7] L. P. S. Singh and C. R. Hagen, "Lagrangian formulation for arbitrary spin. I. the boson case," *Physical Review D*, vol. 9, no. 4, pp. 898–909, 1974.
- [8] F. A. Berends and J. C. J. M. van Reizen, "On spin-3 field theory and the zero-mass limit of higher spin theories," *Nuclear Physics B*, vol. 164, pp. 286–304, 1980.
- [9] C. Aragone, S. Deser, and Z. Yang, "Massive higher spin from dimensional reduction of gauge fields," *Annals of Physics*, vol. 179, no. 1, pp. 76–96, 1987.
- [10] C. C. Chiang, "Lagrangian formalism for neutral, massive spin 3 fields," *Progress of Theoretical Physics*, vol. 45, no. 4, pp. 1311–1320, 1971.
- [11] P. Baikov, M. Hayashi, N. Nelipa, and S. Ostapchenko, "Ghost- and tachyon-free gauge-invariant, Poincaré, affine and projective Lagrangians," *General Relativity and Gravitation*, vol. 24, no. 8, pp. 867–880, 1992.
- [12] M. A. Podoinitsyn, "Polarization spin-tensors in two-spinor formalism and Behrends–Fronsdal spin projection operator for  $D$ -dimensional case," *Physics of Particles and Nuclei Letters*, vol. 16, no. 4, pp. 315–320, 2019.
- [13] M. Kawasaki, M. Kobayashi, and Y. Mori, "Lagrange formulation of the 20-component theory of spin-3 fields," *Lettere al Nuovo Cimento*, vol. 14, no. 17, pp. 611–614, 1975.
- [14] M. Kawasaki and M. Kobayashi, "Lagrange formulation of the symmetric theory of massive spin-3 fields," *Physical Review D*, vol. 17, no. 2, pp. 446–456, 1978.
- [15] D. Dalmazi and E. L. Mendonça, "Higher derivative massive spin-3 models in  $D=2+1$ ," *Physical Review D*, vol. 94, no. 2, article 025033, 2016.
- [16] D. Dalmazi, A. L. R. dos Santos, and R. R. Lino dos Santos, "Higher order self-dual models for spin-3 particles in  $D=2+1$ ," *Physical Review D*, vol. 98, no. 10, article 105002, 2018.
- [17] C. Aragone and A. Khoudeir, "Self-dual spin-4 and 3 theories," 1993, <http://arxiv.org/abs/9307004v1>.
- [18] E. A. Bergshoeff, O. Hohm, and P. K. Townsend, "Massive gravity in three dimensions," *Physical Review Letters*, vol. 102, no. 20, pp. 201–301, 2009.
- [19] H. van Dam and M. Veltman, "Massive and mass-less Yang-Mills and gravitational fields," *Nuclear Physics B*, vol. 22, no. 2, pp. 397–411, 1970.
- [20] A. Leonard, "Aspects of higher spin Hamiltonian dynamics: conformal geometry, duality and charges," PhD thesis, <http://arxiv.org/abs/1709.00719v2>.
- [21] D. Benndorf, D. Dalmazi, and A. L. R. dos Santos, "Hamiltonian positivity of massive spin-2 particles via a rank-2 tensor," *Classical and Quantum Gravity*, vol. 34, no. 4, article 045008, p. 16, 2017.